Numerical simulations of accretion disks - COURSE SYLLABUS





1.	Course title:		
	Numerical simulations of accretion discs		
2.	Lecturer:		
	Miljenko Čemeljić, Bhupendra Mishra		
v3.	Field, type and level of studies, year of study:		
	astrophysics, full-time doctoral studies, all years		
4.	Course character:		
	monographic lecture		
5.	Teaching method:		
	Hybrid, MC in CAMK, BM online		
6.	Language:	English	
7.	Course type and number of hours:		
	lecture 30h		
8.	Estimated load of student's independent work:	24h	
9.	Total workload and number of ECTS points:	54h, 3 ECTS	
10.	Short description and main focus of the course:		
	Accretion processes are among basic mechanisms in the universe. The similarity of underlying principles of accretion at different scales and energetic ranges offer a deep insight with the least assumptions. To obtain anything more than simplest solutions, numerical simulations must be employed. In this course we will outline from the simple ideal hydrodynamical, to most complicated and novel, magnetized accretion disc solut with viscosity and electrical resistivity. We will detail the used numerical methods, and guide the participants through their own simulations. The course will encompass both Newtonian and general relativistic numerical simulations, with hands-on experience the running the codes and visualizing the results in the exercises which will run in parralel the lectures. The acquired knowledge could easily be used in other fields which benef numerical simulations.		
	We will cover the following: -basics of the non-relativistic thin accretion disc theory -analytical solutions for a purely hydrodynamical disc -numerical simulations of a thin disk in purely hydrodynamical setup with the code PLUTO -visualization of the results in 2 and 3 dimensions -simulations of the astrophysical jets and outflows -basics of the general relativistic thin accretion disc theory -hydrodynamical and magnetic instabilities, turbulence		

	-processing of the obtained results for further modelling -basics of General Relativity -pseudopotentials for simulations of black holes, neutron stars and naked singularities -general realtivistic numerical simulations of thin accreton disc		
	The lectures will be shaped in "hands-on" approach: In parallel with the lectures, we will introduce the participants from the very beginning to the numerical simulations. The exercise part of the lectures will be dedicated to this. Our tools will be state-of-the-art codes: a versatile open source code PLUTO for the Newtonian simulations and Athena++ for the general relativistic simulations.		
11.	References:		
	 Books: Frank, King & Raine, "Accretion power in astrophysics" Kato, Fukue & Mineshige, "Black-Hole Accretion Disks: Towards a New Paradigm", Journal papers: Shakura & Sunyaev, 1973, non-relativistic disk Novikov & Thorne, 1973, general relativistic disk M. Čemeljić, 2019, "Atlas" of numerical solutions for star-disk magnetospheric interaction, A&A, 624, A31 Čemeljić, Klużniak, Parthasarathy, 2023, "Magnetically threaded accretion disks in resistive magnetohydrodynamic simulations and asymptotic expansion", A&A 678, A57 Zhu & Stone, 2018, "Global Evolution of an Accretion Disk with a Net Vertical Field: Coronal Accretion, Flux Transport, and Disk Winds", ApJ, 867, 34 B. Mishra, 2016, ."Strongly magnetized accretion discs: structure and accretion from global magnetohydrodynamic simulation", MNRAS 492, 1855 On-line resources: Abramowicz & Straub, Accretion discs, Scholarpedia article: http://www.scholarpedia.org/article/Accretion_discs 		
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12. 13.	http://www.scholarpedia.org/article/Accretion_discs Prerequisites: General physics level of mechanics, heat and thermodynamics basic level of General Relativity theory, basic usage of linux Educational outcomes:	s and electromagnetism, PQF level 8 codes:	
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	At least 70% attendance, final grade depends on the evaluation of the report.	
16.	Contact with the lecturer:	
	Email: miki,mbhupe@camk.edu.pl, consultation in CAMK, office A14 with M.Cemeljic, online with B.Mishra	